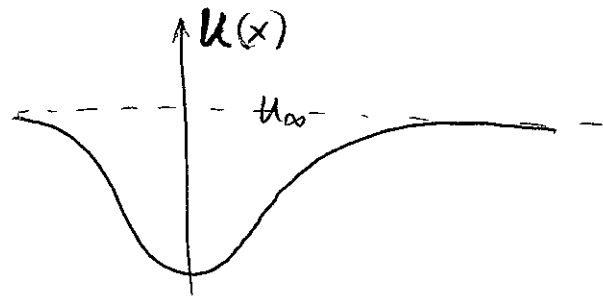


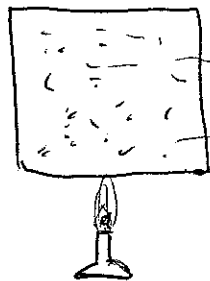
LAST TIME :



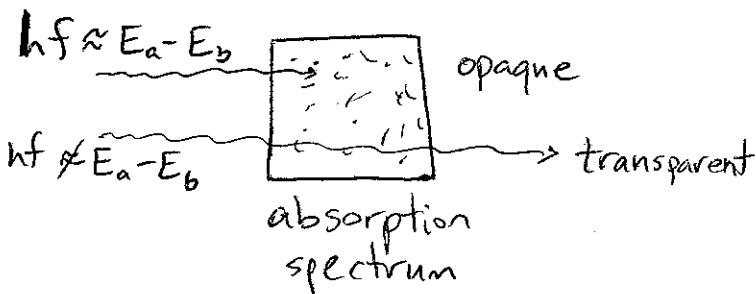
- states with $E < U_{\infty}$ exist only for specific energies

these energies determine most observable properties of the system.

e.g.
emission spectrum

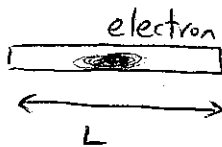


$hf = E_a - E_b$
different for different atoms/molecules

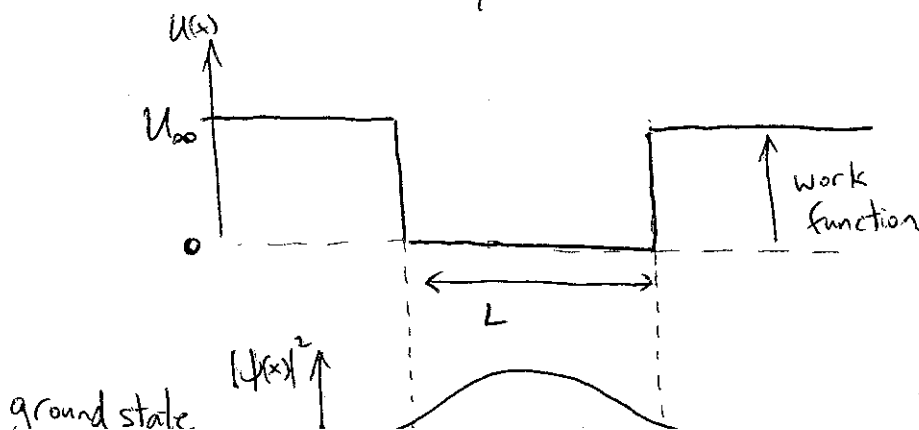


conductor: free electrons
- all possible energies
- opaque for all wavelengths.

Another example: thin short wire



model by "finite square well"



- Number and energy of bound states determined by L, U_0
- Ground state energy > 0
 - * quantum electron always has some kinetic energy *
 - ZERO POINT ENERGY
- Wavefunction extends into region with $E < U(x)$
 - some probability of finding electron in region where it doesn't have enough energy to get classically.

dramatic consequence: TUNNELING



electron with $E < U_0$ can "jump" from one wire to other.

(SIM)

CLICKER

Energy eigenstates allow us to find time dependence of general wavefn:

$$\psi(x, t=0) = \sum c_n f_n(x) \Rightarrow \psi(x, t) = \sum c_n f_n(x) e^{-\frac{iE_n t}{\hbar}}$$


↑
 energy eigenstate.

prob. dens. → const for energy eigenstate
 → not for sum of eigenstates.

What about states with $E > U_\infty$?

- all energies $E > U_\infty$ allowed

- BUT: not normalizable unless we form wavepackets.

e.g. electron wavepacket 

•
PROTON

evolution of
wavepacket describes
SCATTERING from proton

SIM

