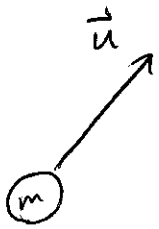


LAST TIME:



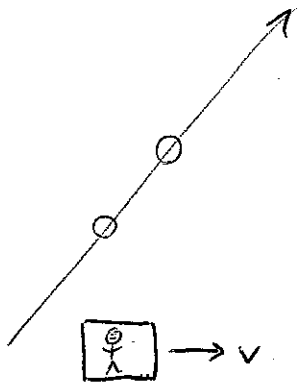
Momentum: $\boxed{\vec{p} = \gamma_{\vec{u}} m \vec{u}} = m \frac{d\vec{x}}{d\tau}$

in any frame, sum of \vec{p} for all objects same before & after collision.

TODAY: Energy: $\boxed{E = \gamma m c^2} = m c^2 \frac{dt}{d\tau}$

Lorentz tforms mix \vec{x} and t

\therefore will mix $m \frac{d\vec{x}}{d\tau}$ with $m \frac{dt}{d\tau}$



$$dz' = dz$$

$$dy' = dy$$

$$dx' = \gamma_v (dx - v dt)$$

$$dt' = \gamma_v (dt - \frac{v}{c^2} dx)$$

$$\xrightarrow{\times m \cdot \frac{1}{d\tau}}$$

$$p'_z = p_z$$

$$p'_y = p_y$$

$$p'_x = \gamma_v (p_x - v (\frac{E}{c^2}))$$

$$E' = \gamma_v (E - v p_x)$$

$$p'_x = \gamma_v (p_x - \frac{v}{c^2} E)$$

\uparrow conserved \uparrow conserved \uparrow must be conserved.

Small \vec{u} :

$$E = \frac{m c^2}{\sqrt{1 - \frac{u^2}{c^2}}} \approx m c^2 + \frac{1}{2} m u^2 + \dots$$

\uparrow mass $\times c^2$ \uparrow kinetic energy

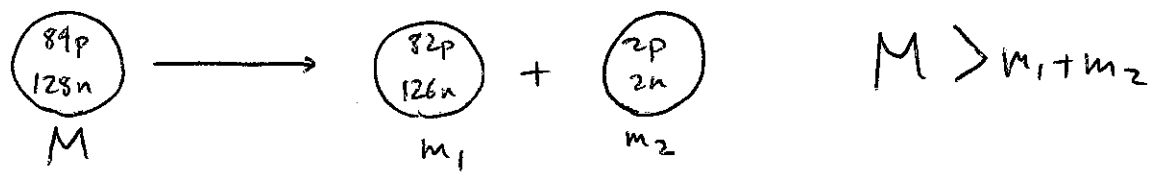
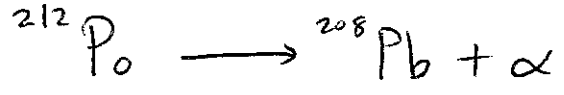
classical: - mass always conserved
 - kinetic energy only conserved in elastic collisions.

relativity: $E = mc^2 + (\gamma - 1)mc^2$ Conserved in ALL processes.

\uparrow MASS ENERGY \uparrow RELATIVISTIC KINETIC ENERGY

BUT: can convert mass \longleftrightarrow kinetic energy CLICKER.

e.g. nuclear reactions



$(\Delta M)c^2 \approx 1.4 \times 10^{-12} \text{ J} \rightarrow$ goes into kinetic energy of products.

\rightarrow enormous if multiplied by Avogadro's #.

Analyze via cons. of energy + momentum:

BEFORE:



Momentum cons:

$$0 = -\gamma_1 m_1 v_1 + \gamma_2 m_2 v_2$$

AFTER:

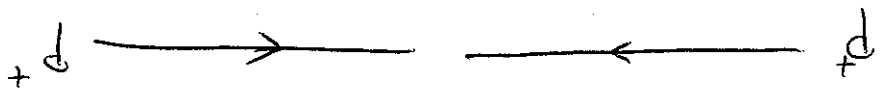


Energy cons:

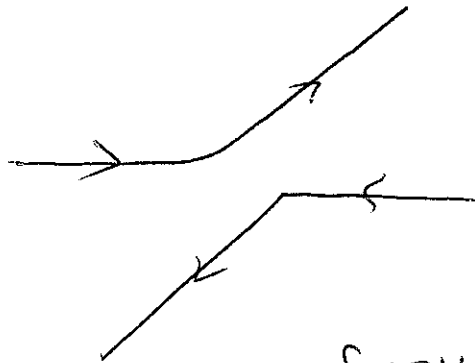
$$Mc^2 = \gamma_1 m_1 c^2 + \gamma_2 m_2 c^2$$

- can solve for v_1 + v_2

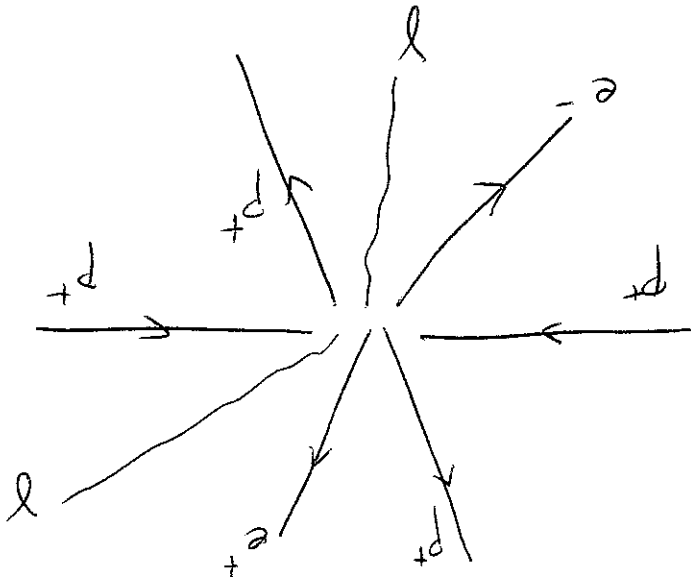
Another example: particle collisions



low E: scattering



high E:



can produce new particles from kinetic energy!

LHC: each proton has $E \approx 7000$ times its rest energy

→ could produce 14000 protons/anti-protons.