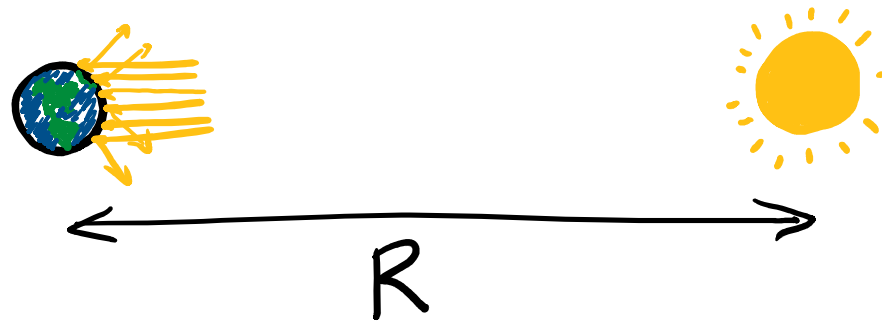


Learning goals

- Explain quantitatively how the intensity of light from an object radiating uniformly in all directions varies with the distance to the object
- Describe which molecular properties of a gas affect the pressure on the walls of a container, and what proportionality relationship each of these quantities has with the pressure
- Describe the microscopic origin of the ideal gas law

Office hours: today 3-5pm, Tuesday 3-4:30pm

Hennings 420



If we moved the Earth twice as far away from the Sun, the power of solar radiation hitting the Earth would be

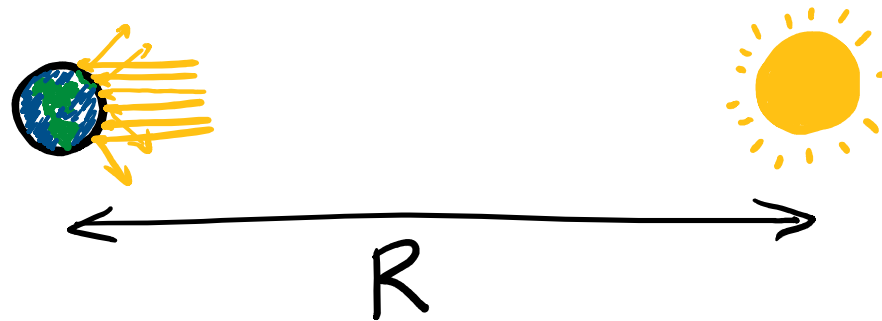
- A) twice as much as before.
- B) the same as before.
- C) half as much as before
- D) one quarter as much as before.
- E) one eighth as much as before.

No tutorials
this week.

Midterm Q&A
Tuesday 5-7
Henn 202

Office hours: today 3-5pm, Tuesday 3-4:30pm

Hennings 420



If we moved the Earth twice as far away from the Sun, the power of solar radiation hitting the Earth would be

- A) twice as much as before.
- B) the same as before.
- C) half as much as before.
- D) one quarter as much as before.**
- E) one eighth as much as before.

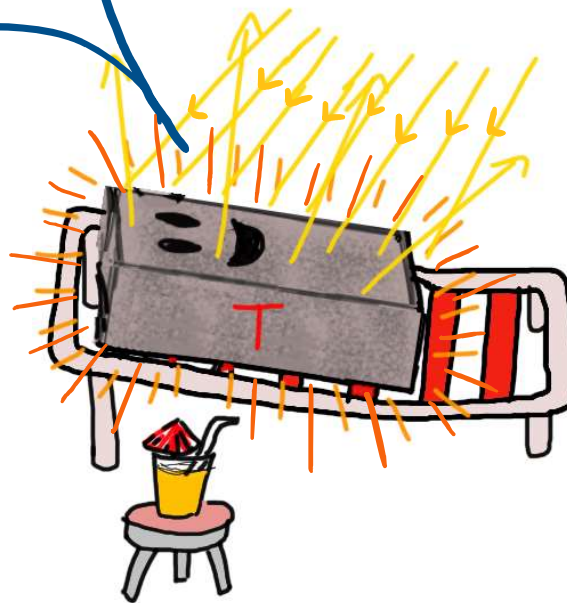


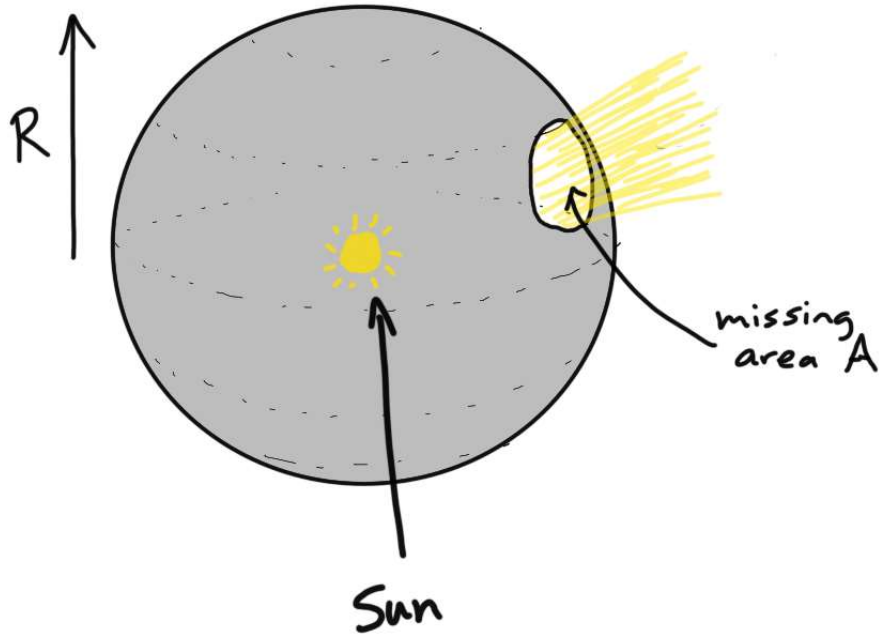
At twice the distance, sunlight spreads over $4 \times$ the area, so Power/Area is $\frac{1}{4}$
↑ intensity

No tutorials this week.

Midterm Q&A
Tuesday 5-7
Henn 202

Last time
in Physics 157...





Power through hole is

$$H_{sun} \cdot \frac{A}{4\pi R^2}$$

INTENSITY of sunlight (power per area) is

$$I = \frac{H_{sun}}{4\pi R^2} \rightarrow \text{double } R \Rightarrow \frac{1}{4} I$$

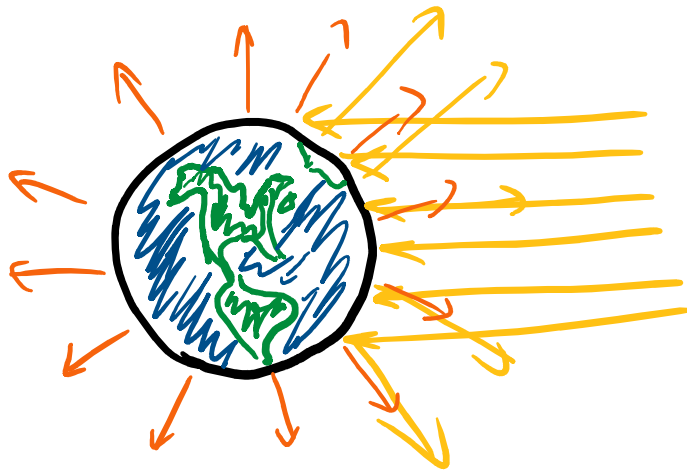
At distance of Earth, $I = I_{sc} = 1367 \text{ W/m}^2$

Key relation for steady-state heat flow:

$$H_{in} = H_{out}$$



Our problem:



★ set equal & solve for T ★

thermal
 H_{out} : IR radiation

$$= (4\pi r^2) \epsilon \sigma T^4$$

↑ $A_{surface}$

H_{in} : absorbed sunlight

$$I_{sc} \cdot \pi r^2 \cdot (1 - a)$$

↑
albedo = fraction reflected

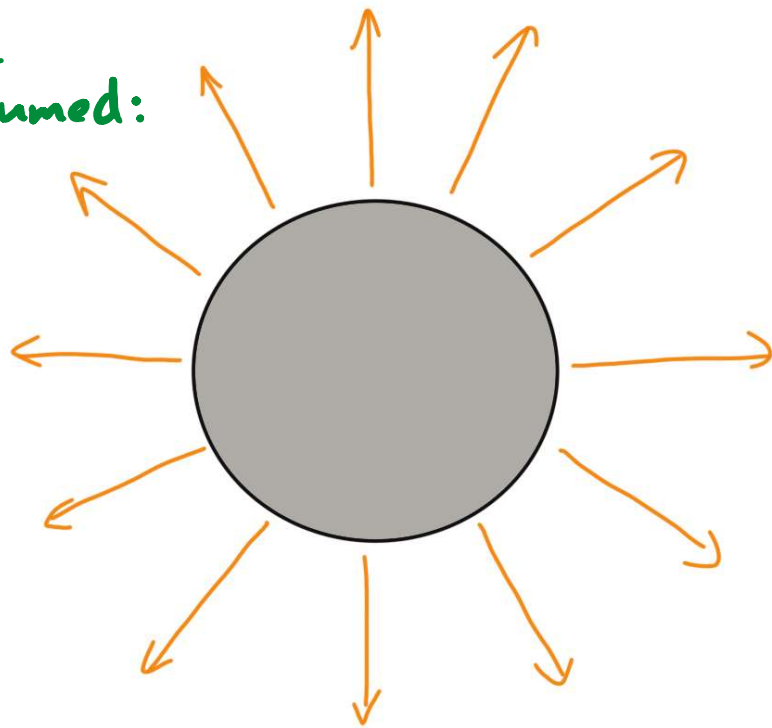
Result:

$$T = \left(\frac{I_{sc} \cdot (1-a)}{4 e \sigma} \right)^{\frac{1}{4}}$$

Gives $T = -18^{\circ}\text{C}$ for $e=1$ but...

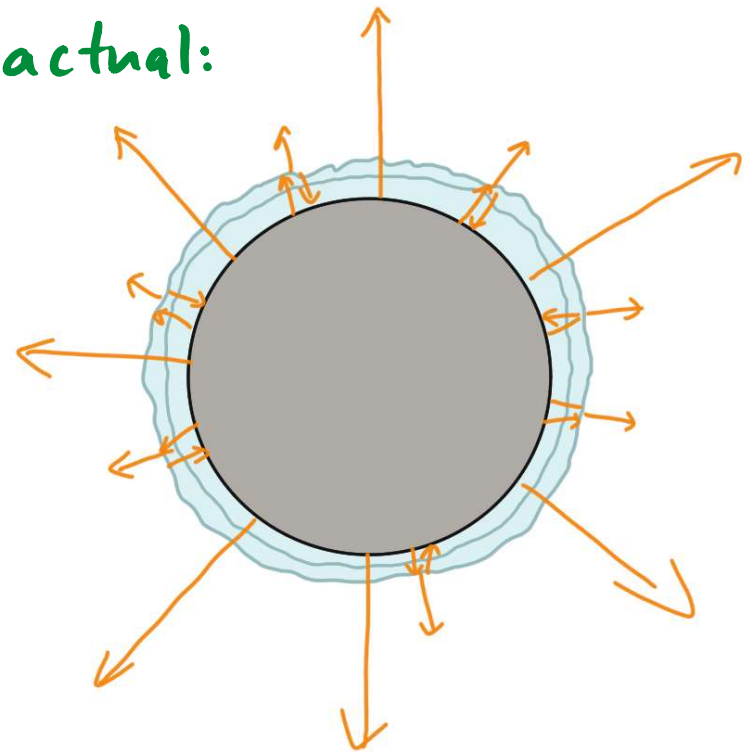
Actual surface temperature is larger due to the
GREENHOUSE EFFECT: some IR radiation is
absorbed by "greenhouse gases" + re-emitted back to
Earth.

we
assumed:



$e = 1$

actual:



$e \approx 0.6$

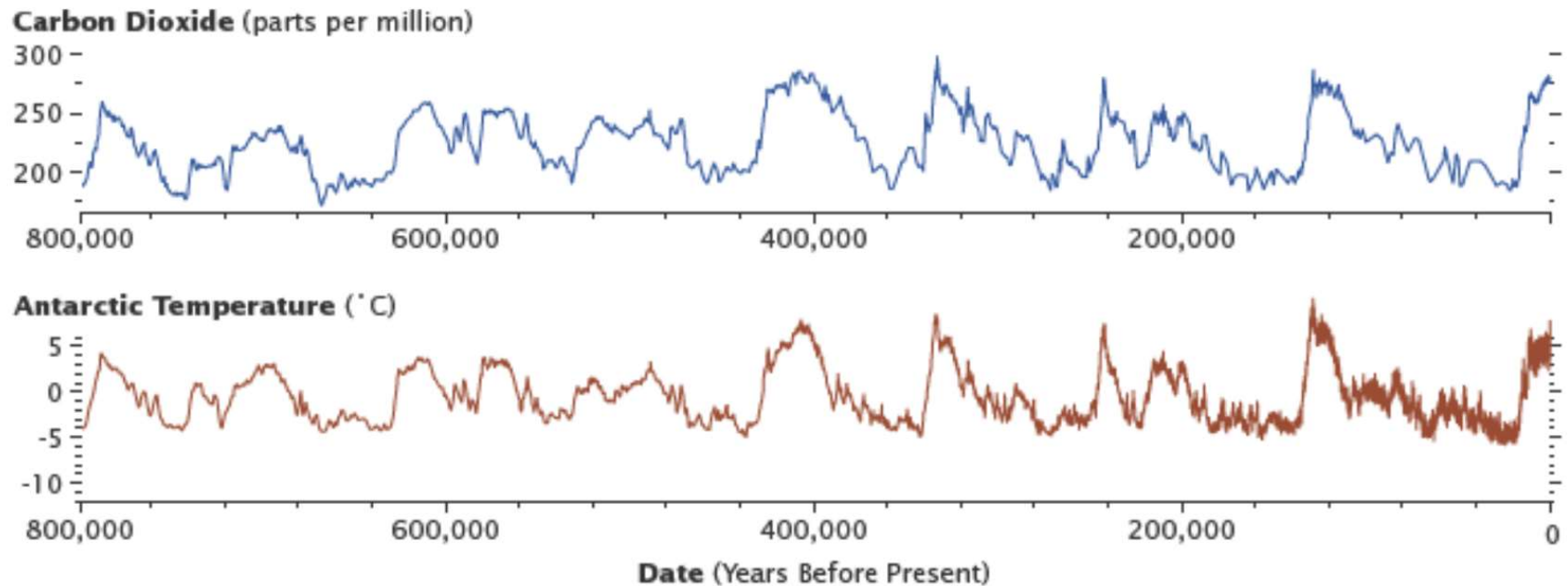
$$\star T = \left[\frac{(1-a)I_{sc}}{4e\sigma} \right]^{\frac{1}{4}} \star$$

Lower $e \Rightarrow$ higher T

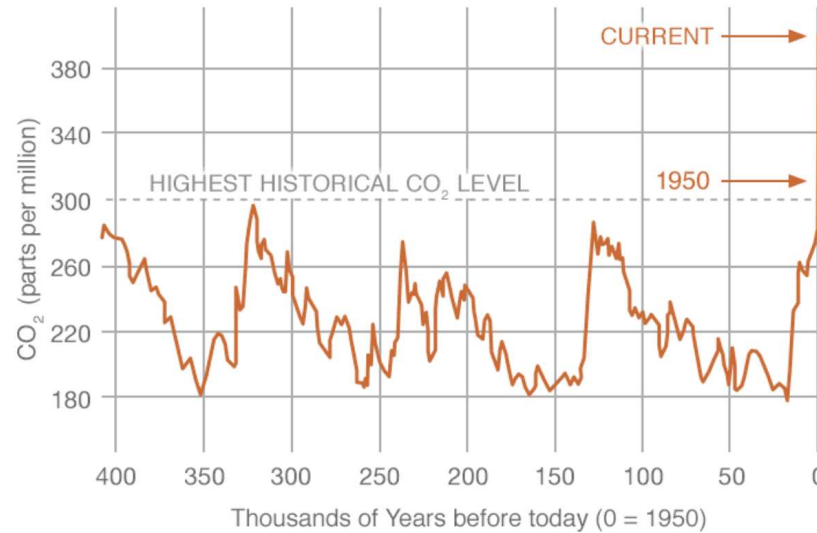
Real $e \approx 0.6$ gives $T = 14.5^\circ\text{C}$

But e can decrease e.g. due to increasing CO_2 concentration in atmosphere. \longrightarrow Global warming

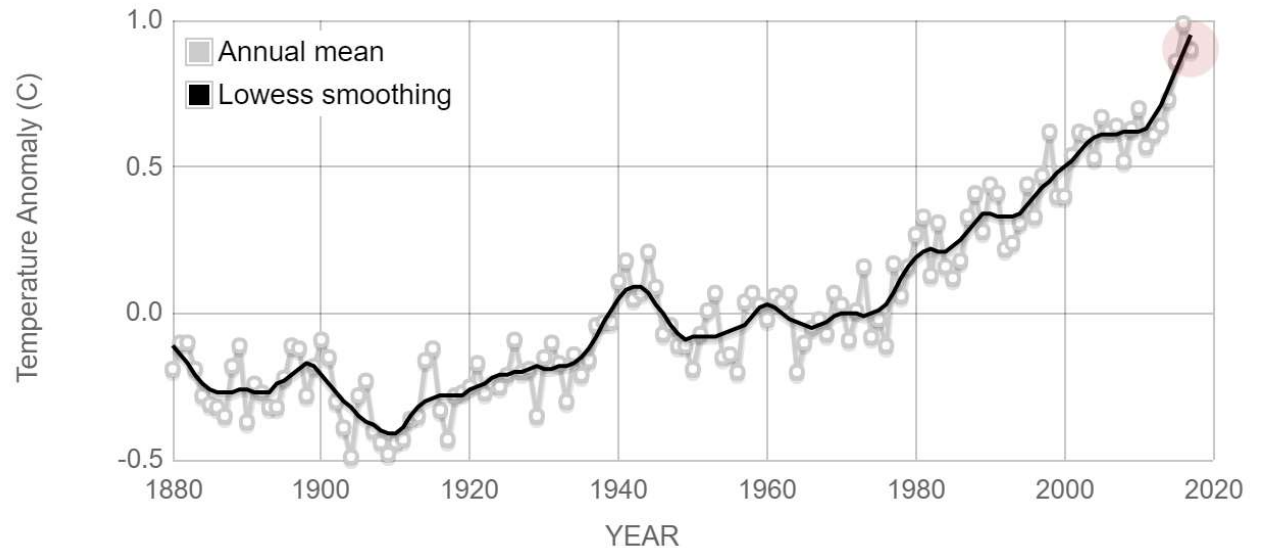
CO₂ correlates closely with temperature



CO₂ levels:



Temperature:



Almost all climate scientists believe this rise due to human activity

2. Mars albedo, the reflection coefficient for sunlight from Mars, is 0.250. The radius of Mars is 3397 km. The Solar constant at Earth is 1367 W/m^2 and the distance from Mars to the Sun is 1.52 times the Earth to Sun distance.

a) Find the temperature of Mars.

power of
absorbed sunlight



Q: Write an expression for H_{in}
in terms of the information provided

(you don't need to evaluate it)



2. Mars albedo, the reflection coefficient for sunlight from Mars, is 0.250. The radius of Mars is 3397 km. The Solar constant at Earth is 1367 W/m^2 and the distance from Mars to the Sun is 1.52 times the Earth to Sun distance.

a) Find the temperature of Mars.

power of
absorbed sunlight



Q: Write an expression for H_{in}
in terms of the information provided

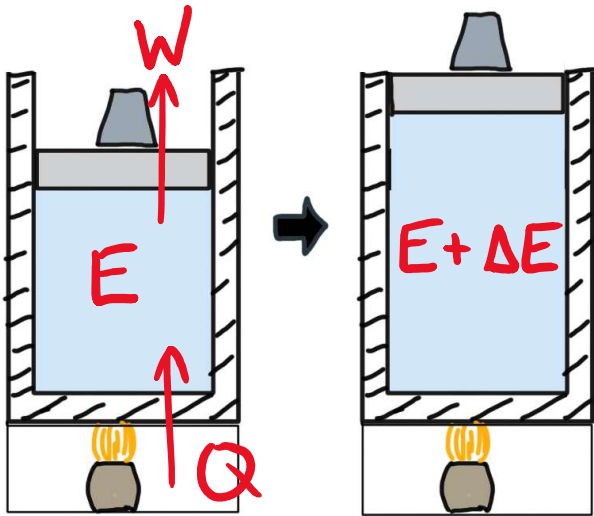
(you don't need to evaluate it)

for Earth $H_{in} = I_{sc} \times \pi r_E^2 \times (1 - a_E)$

for Mars: $H_{in} = I_{Mars} \times \pi r_M^2 \times (1 - a_M)$

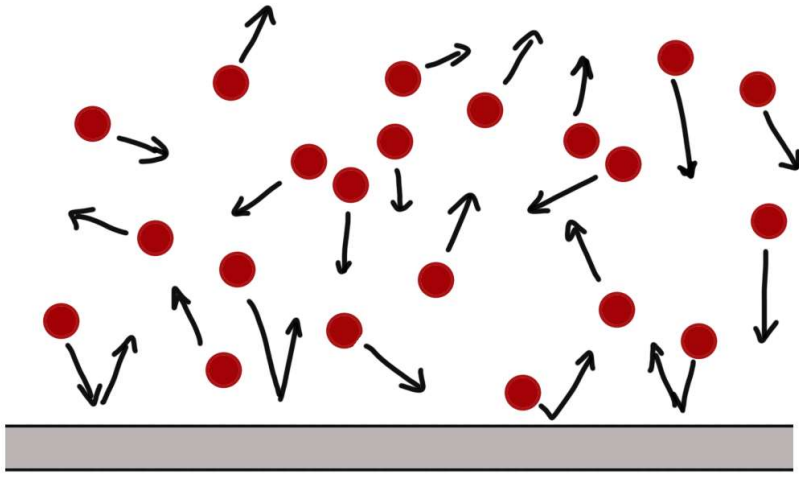
$$R_{Mars} = 1.52 R_{Earth} \text{ so } I_{Mars} = \frac{1}{1.52^2} I_{Earth}$$

New Topic: THE FIRST LAW OF THERMODYNAMICS
= Conservation of energy



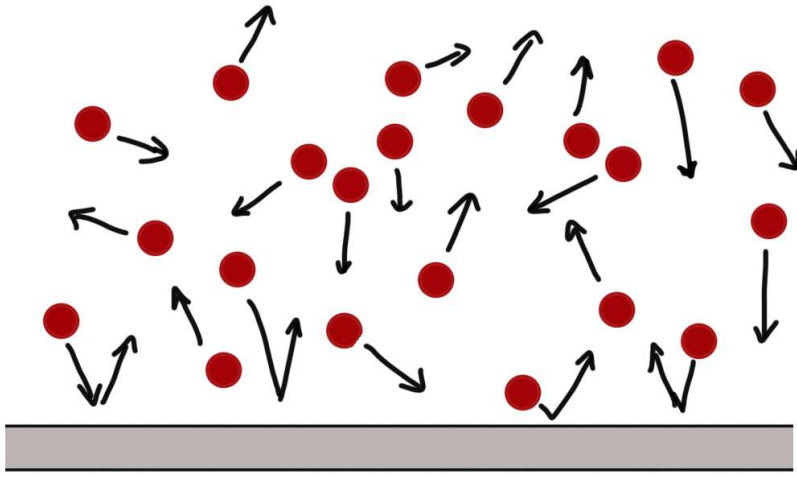
$$\Delta E_{\text{gas}} = Q - W$$

work done
by gas



The picture shows molecules of an ideal gas near the wall of a container. What properties of these molecules does the pressure on the wall (force per unit area) depend on?

EXTRA: for each quantity you identify, what would happen to the pressure if you double that quantity?



The picture shows molecules of an ideal gas near the wall of a container. What properties of these molecules does the pressure on the wall (force per unit area) depend on?

double density \rightarrow double P

double mass \rightarrow double P

double velocity \rightarrow quadruple P

(twice as many collisions
twice as much impact)

$$P = \text{const.} \cdot \frac{N}{V} \cdot \overbrace{m \cdot v_{\text{avg}}^2}^{\substack{\text{2x kinetic energy} \\ \text{per molecule}}}$$

density \downarrow

TEMPERATURE & KINETIC ENERGY

If we define $T = \text{constant}_1 \times E_{\text{kin}}^{\text{avg}}$

Molecular model gives:

$$P = \text{constant}_2 \times \frac{N}{V} \times T$$

So P is proportional to T for fixed n, V

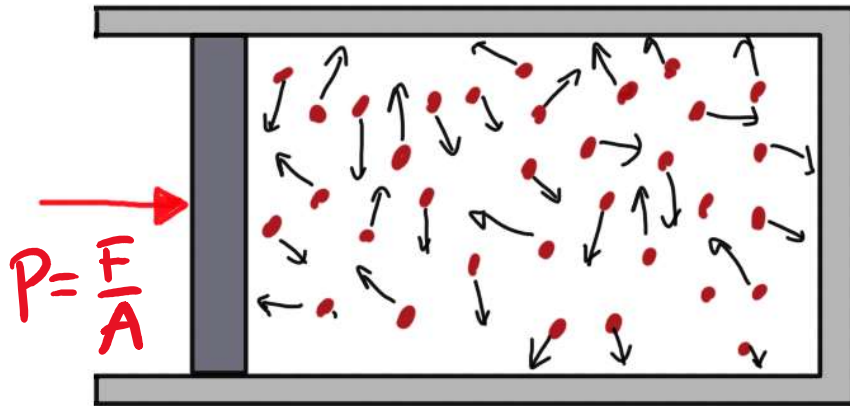
* Molecular definition of temperature is consistent with Kelvin scale definition!

Definitions match exactly if:

$$\text{constant}_2 = \frac{8.31 \text{ J/mol}\cdot\text{K}}{6.02 \times 10^{23}} \leftarrow \text{call this } R$$

\leftarrow this is N_A

IDEAL GAS LAW



molecular density

avg. kinetic energy per molecule

$$P = \text{const} \cdot \frac{N}{V} \cdot E_{\text{avg}}^{\text{kin}}$$



$$PV = n \cdot R \cdot T$$

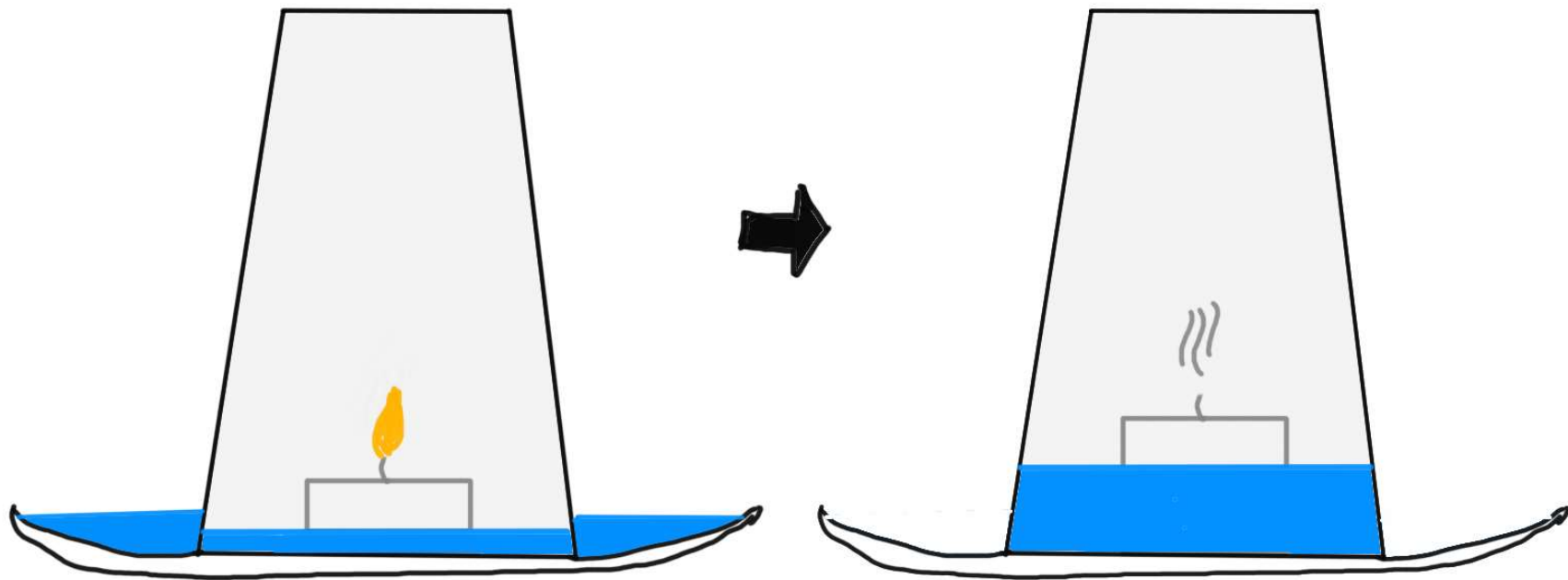
moles

$8.31 \frac{\text{J}}{\text{mol} \cdot \text{K}}$

Tells us how much force a gas exerts on the wall

Thirsty cup demo

https://www.youtube.com/watch?v=3EGfqU_zBec



Why?

Hint: $PV = nRT$